

#### Groundbreaking Ceremony for the SuperKEKB Project, KEK, November 18, 2011



## The SuperKEKB Project

#### Peter Križan University of Ljubljana and J. Stefan Institute





#### Thank you for coming, in spite of your very busy schedules!

#### Contents

•Scope of the project

●Accelerator upgrade → SuperKEKB

•Detector upgrade → Belle-II

•Status and outlook





Peter Križan, Ljubljana

#### A little bit of history...

**CP violation:** difference in the properties of particles and their anti-particles – first observed in 1964.

**M. Kobayashi and T. Maskawa (1973):** CP violation in the Standard model – related to the weak interaction quark transition matrix

Their theory was formulated at a time when three quarks were known – and they requested the existence of three more!

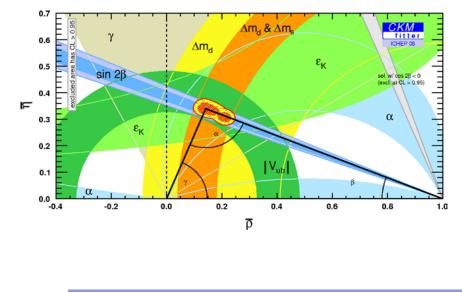
The last missing quark was found in 1994.

... and in 2001 two experiments – Belle and BaBar at two powerfull accelerators (B factories) - have further investigated CP violation and have indeed proven that it is tightly connected to the quark transition matrix

## KM's bold idea verified by experiment







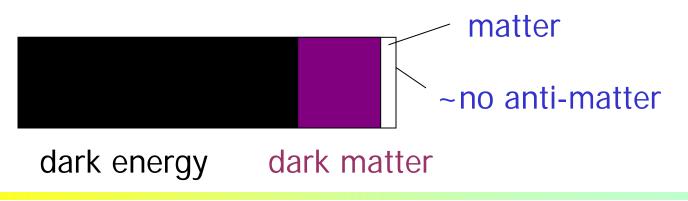


# → With essential experimental confirmations by Belle and BaBar! (explicitly noted in the Nobel Prize citation)

# The KM scheme is now part of the Standard Model of Particle Physics

•However, the CP violation of the KM mechanism is too small to account for the <u>asymmetry between matter and anti-matter</u> in the Universe (falls short by 10 orders of magnitude !)

- •SM does not contain the fourth fundamental interaction, gravitation
- •Most of the Universe is made of stuff we do not understand...



# Are we done? (Didn't the B factories accomplish their mission, recognized by the 2008 Nobel Prize in Physics?)





И зарачения С. Окубо при больтой петерабура аля Вселенной стица науба но се кривой аринуте нарушение ср-инвариантности, с-асниметрия и Барионная асимиетрия вселенной

A.A.Cazapos

Теория расширяющейся Бселенной, предполагающая свёрхалотное начальное состояние вещества, по-видимому, исключает возможность макроскопического разделения вещества и антивещества; поэтому следует Matter - anti-matter asymmetry of the Universe: KM (Kobayashi-Maskawa) mechanism still short by 10 orders of magnitude !!! Two complementary approaches to study shortcomings of the Standard Model and to search for the so far unobserved processes and particles (so called New Physics, NP). These are the **energy frontier** and the **intensity frontier**.

**Energy frontier :** direct search for production of unknown particles at the highest achievable energies.

**Intensity frontier :** search for rare processes, deviations between theory predictions and experiments with the ultimate precision.

→for this kind of studies, one has to investigate a very large number of reactions ("events") → need accelerators with ultimate intensity ("luminosity")

#### Comparison of energy /intensity frontiers

To observe a large ship far away one can either use **strong binoculars** or observe **carefully the direction and the speed of waves** produced by the vessel.

#### **Energy frontier (LHC)**

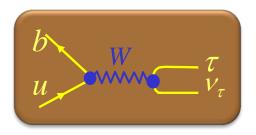




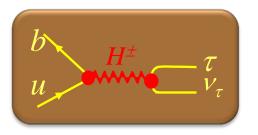


#### An example: Hunting the charged Higgs in the decay $B^- \rightarrow \tau^- \nu_{\tau}$

In addition to the Standard Model Higgs to be discovered at the LHC, in New Physics (e.g., in supersymmetric theories) there could be another 'God particle' – a charged Higgs.



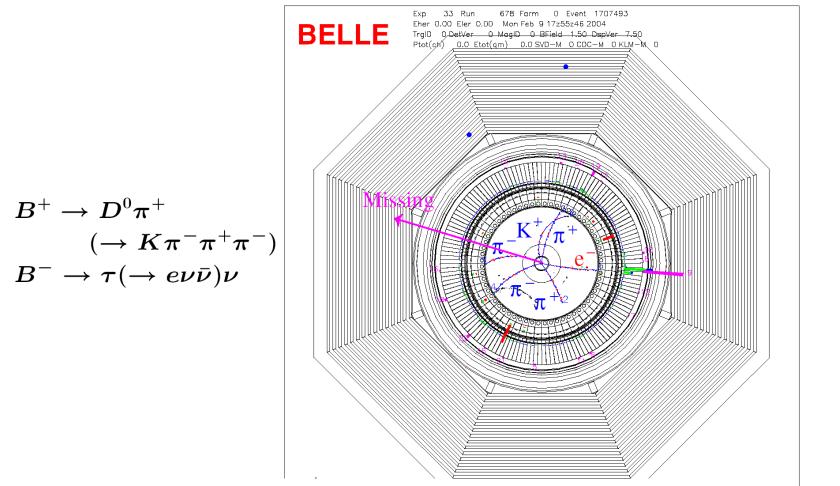
The rare decay  ${\rm B}^{\scriptscriptstyle -} \to \tau^{\scriptscriptstyle -}\,\nu_\tau$  is in SM mediated by the W boson



In some supersymmetric extension it can also proceed via a charged Higgs

The charged Higgs would influence the decay of a B meson to a tau lepton and its neutrino, and modify the probability for this decay.

#### Missing Energy Decays: $B^{-} \rightarrow \tau^{-} \nu_{\tau}$

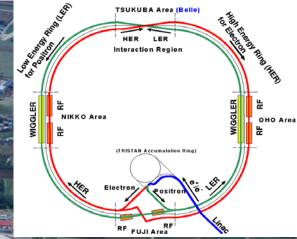


By measured the decay probability (branching fraction) and comparing it to the SM expectation:

→ Properties of the charged Higgs (e.g. its mass)

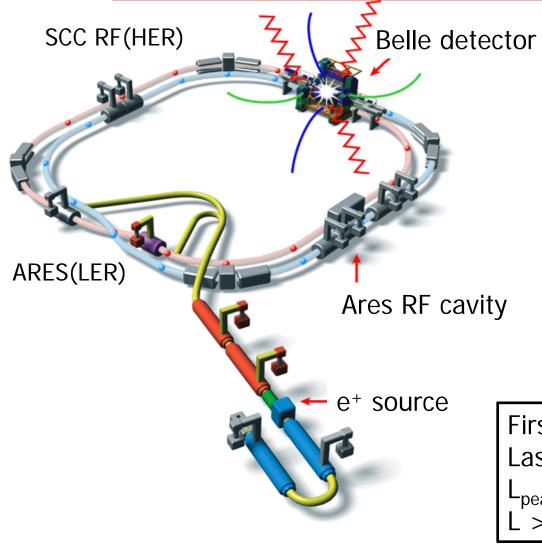
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## 



## The KEKB Collider

Fantastic performance far beyond design values!



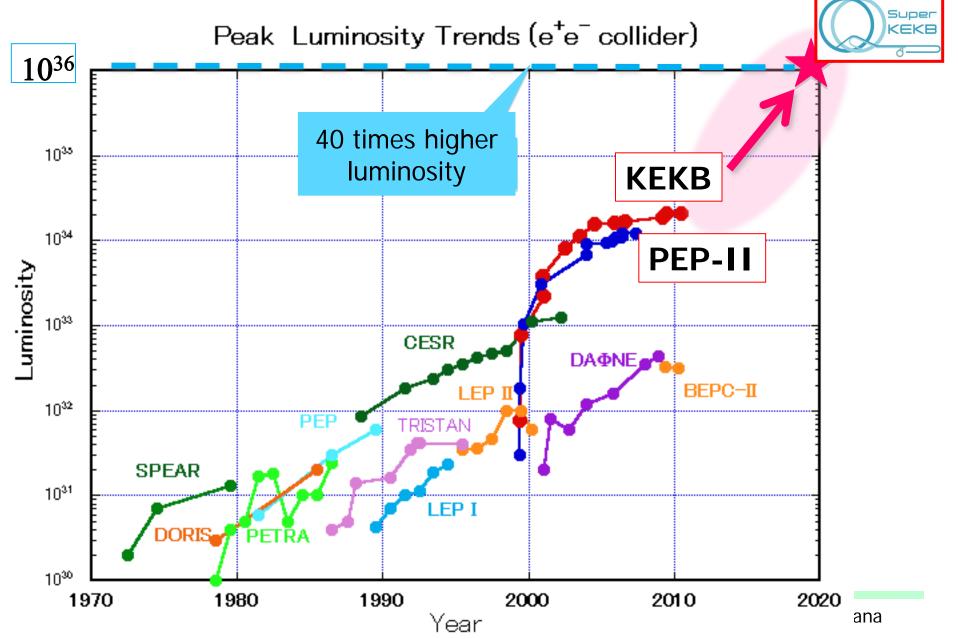
- e<sup>-</sup> (8 GeV) on e<sup>+</sup>(3.5 GeV)

- √s ≈ m<sub>γ(4S)</sub>
- Lorentz boost:  $\beta \gamma = 0.425$
- 22 mrad crossing angle

Peak luminosity (WR!) : **2.1 x 10<sup>34</sup> cm<sup>-2</sup>s<sup>-1</sup>** =2x design value

First physics run on June 2, 1999 Last physics run on June 30, 2010  $L_{peak} = 2.1x10^{34}/cm^2/s$ L > 1ab<sup>-1</sup>

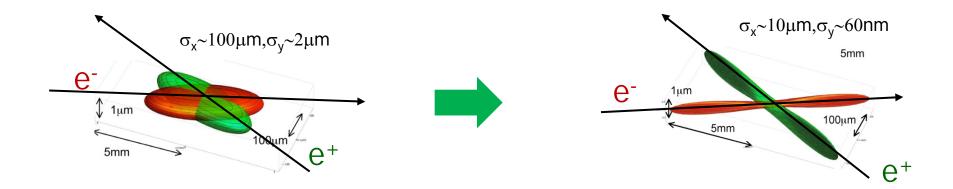
## SuperKEKB is the intensity frontier





How to go from an excellent accelerator with world record performance – KEKB – to a 40x times better, more intense facility?

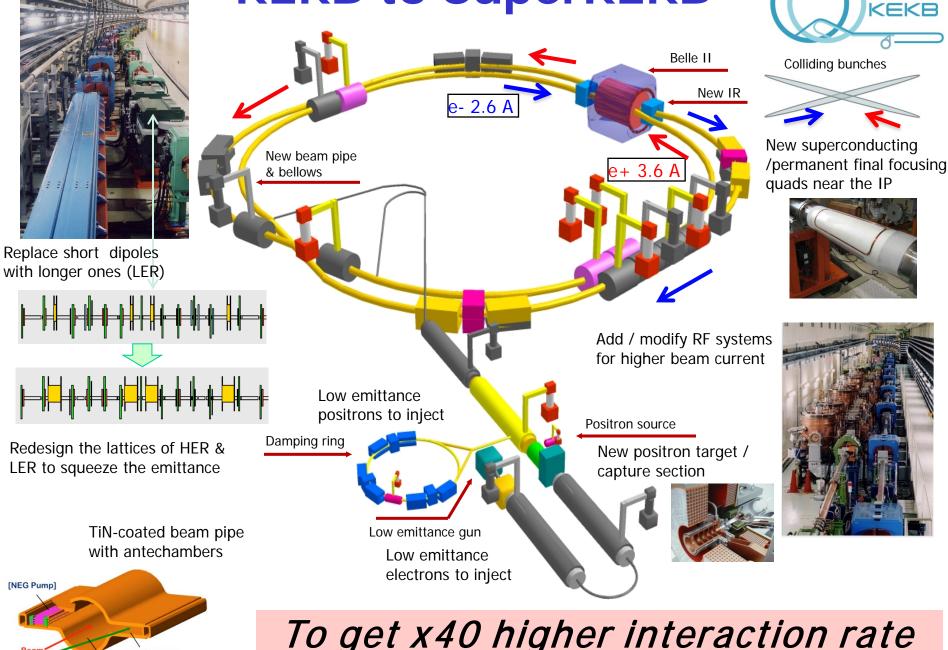
In KEKB, colliding electron and positron beams are much thinner than the human hair...



... For a 40x increase in intensity you have to make the beam as thin as 100 atomic layers!

Peter Križan, Ljubljana





[SR Channel] [Beam Channel]

#### To get x40 higher interaction rate

Super

## Need to build a new detector to handle higher backgrounds

Critical issues at L= 8 x  $10^{35}$ /cm<sup>2</sup>/sec

- Higher background ( ×10-20)
  - radiation damage and occupancy
  - fake hits and pile-up noise in the EM
- Higher event rate ( ×10)
  - higher rate trigger, DAQ and computing
- Require special features
  - low  $p \mu$  identification  $\leftarrow$  s $\mu\mu$  recon. eff.
  - hermeticity  $\leftarrow v$  "reconstruction"

Have to employ and develop very advanced technologies to build such an appartus! BELLE 1 MagID 21 BField 1.50 DspVer 7.50 TrgID 0 DetVer

TDR published arXiv:1011.0352v1 [physics.ins-det]

 $\rightarrow$ 

# Belle II Detector

KL and muon detector: Resistive Plate Counter (barrel) Scintillator + WLSF + MPPC (end-caps)

EM Calorimeter: CsI(TI), waveform sampling (barrel) Pure CsI + waveform sampling (end-caps)

electrons (7GeV)

Beryllium beam pipe 2cm diameter

Vertex Detector 2 layers DEPFET + 4 layers DSSD

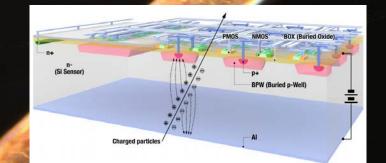
> Central Drift Chamber He(50%):C<sub>2</sub>H<sub>6</sub>(50%), Small cells, long lever arm, fast electronics

Particle Identification Time-of-Propagation counter (barrel) Prox. focusing Aerogel RICH (fwd)

positrons (4GeV)

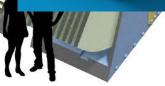
Determine the reaction point position with a fantastic precision - extremly delicate elements

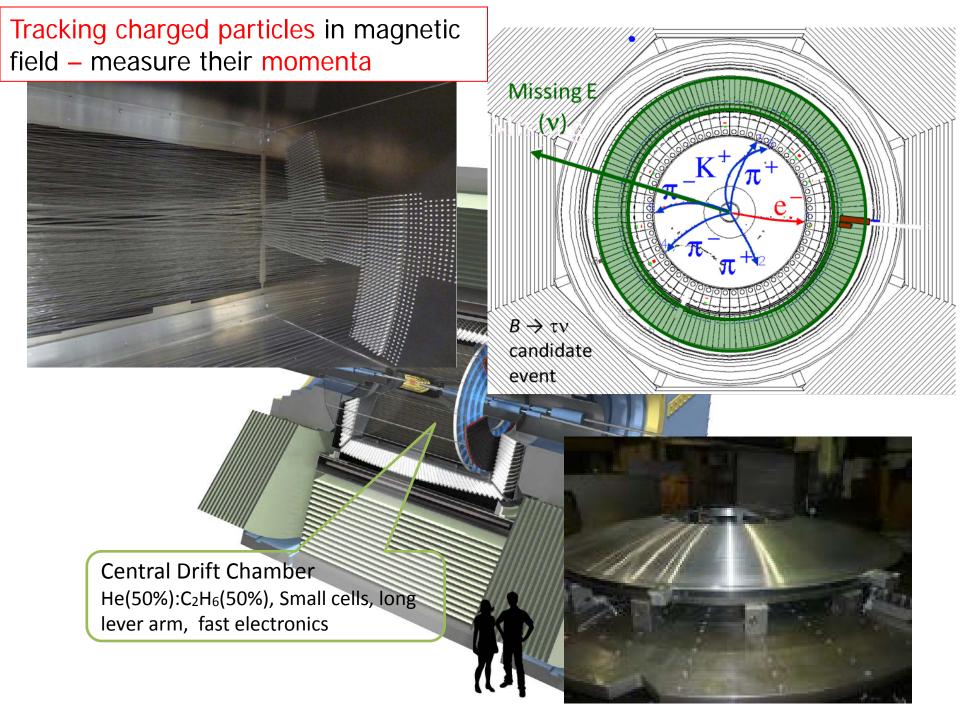
### Hair – 100 microns thick



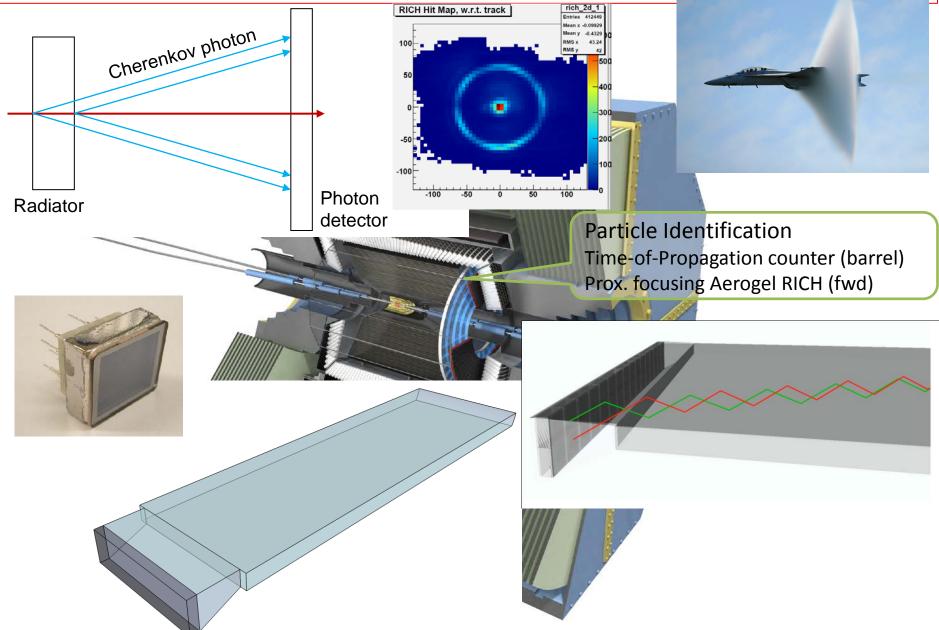
Beryllium beam pipe 2cm diameter

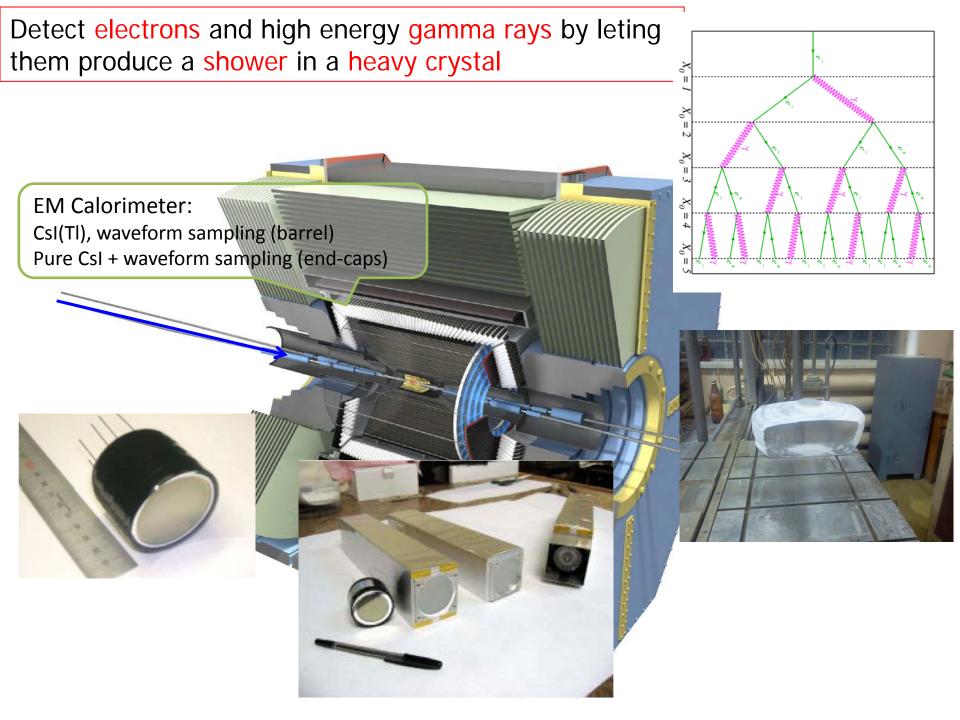
Vertex Detector 2 layers DEPFET + 4 layers DSSD



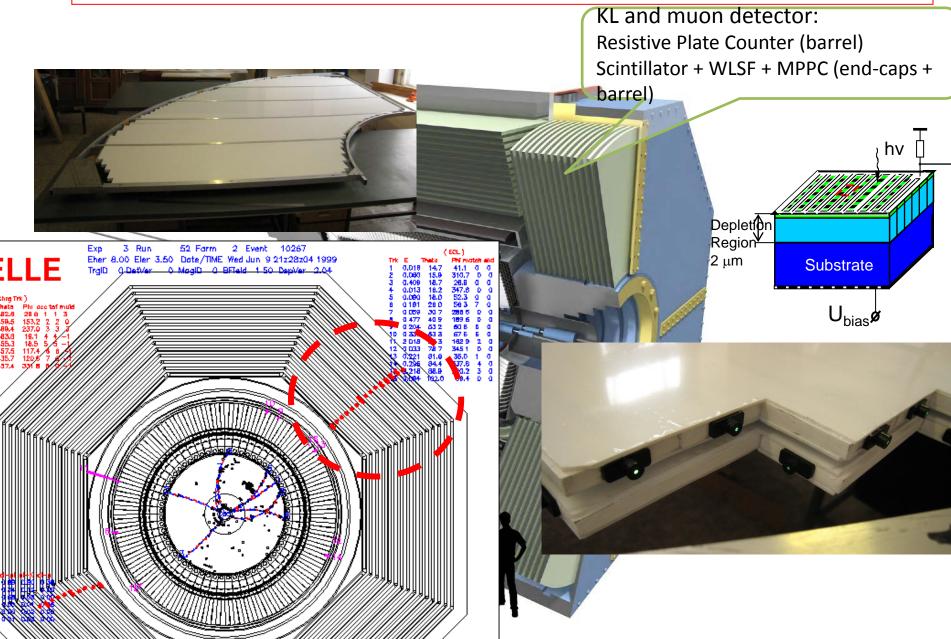


Use Cherenkov effect: light emitted by a particle faster than velocity of light in a medium - like a shock wave from a supersonic airplane!





#### Detect muons: particles that penetrate 1m of iron



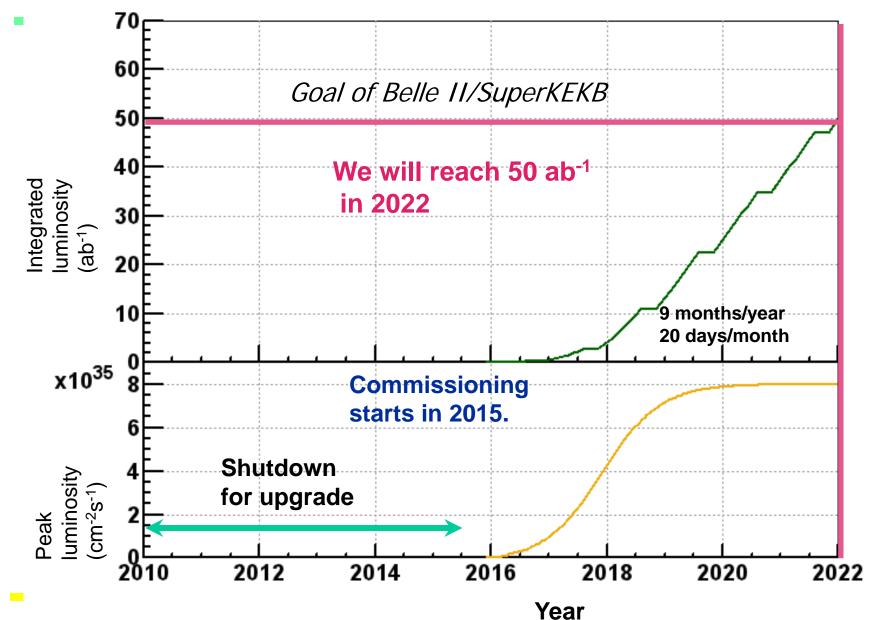
# The Belle II Collaboration



A very strong group of ~400 highly motivated scientists!

# Schedule (Beam starts in Fall 2014)







### Conclusion



- KEKB has proven to be an excellent tool for flavour physics, with reliable long term operation, breaking world records, and surpassing its design perfomance by a factor of two.
- Major upgrade at KEK in 2010-14 → SuperKEKB+Belle II, with 40x larger event rates, construction started
- Expect a new, exciting era of discoveries, complementary to the LHC

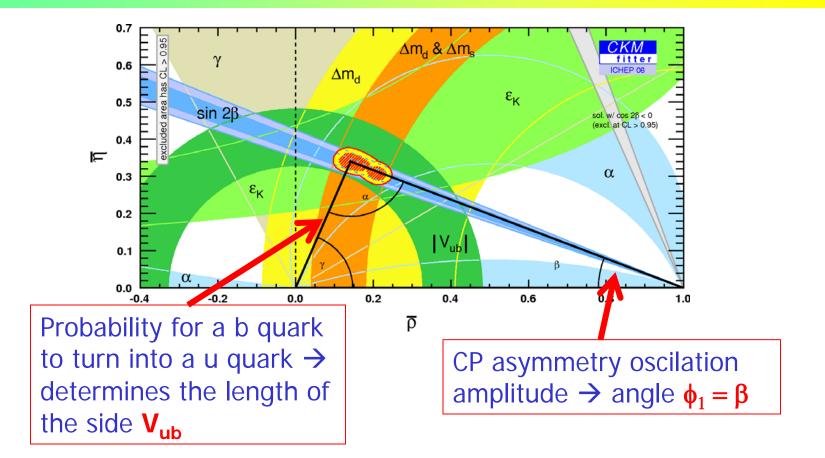
We scientists and experts have come from all over the world to join forces in this exciting project. We are very enthusiastic about it, so let us work together and accomplish it successfully!

このエキサイティングなプロジェクトに参加するために 世界中から優秀な科学者 が集まりました.みな,とても熱くなっています.このプロジェクトの成功のために 共に努力したいと 思います.

#### More slides....

Peter Križan, Ljubljana

#### All experimental studies combined...



Constraints from measurements of angles and sides of the unitarity triangle

→ Remarkable agreement

B Physics @ Y	(4S)						
-	<u>`</u>		-1-1\	$\frac{\text{Observable}}{ V_{cb}  \text{ (exclusive)}}$	$\frac{B \text{ Factories } (2 \text{ ab}^{-1})}{4\% (*)}$	Super $B (75 \text{ ab}^{-1})$ 1.0% (*)	
	. ,	Super $B(7)$		$ V_{cb} $ (exclusive) $ V_{cb} $ (inclusive)	470 (*) 1% (*)	0.5% (*)	
$\sin(2eta) \; (J/\psi^{\cdot}K^{0}) \ \cos(2eta) \; (J/\psi^{\cdot}K^{st 0})$	0.018 0.30	0.005 0.03		$ V_{cb} $ (inclusive) $ V_{ub} $ (exclusive)	170 (*) 8% (*)	0.5% (*) 3.0% (*)	
$\sin(2\beta) (Dh^0)$	0.10	0.0		$ V_{ub} $ (exclusive) $ V_{ub} $ (inclusive)	• /	2.0% (*)	
$ cos(2\beta) (Dh^0) $	0.20	0.04		$ V_{ub} $ (inclusive)	8% (*)	2.0% (*)	
$S(J/\psi \pi^0)$	0.10	0.0		<b>1</b> 2( <b>D</b> )	20%	407 (4)	
$S(D^+D^-)$	0.20	0.0		$\mathcal{B}(B \to \tau \nu)$		4% (†)	
$S(\phi K^0)$	0.13	0.02		$\mathcal{B}(B \to \mu \nu)$	visible	5%	
$S(\eta'K^0)$	0.05	0.01 (	(*)	$\mathcal{B}(B  o D  au  u)$	10%	2%	
$S(K_s^0K_s^0K_s^0)$	0.15	0.02 (				207 (V)	
$S(K_g^0\pi^0)$	0.15	0.02 (	(*)	$\mathcal{B}(B  o  ho \gamma)$	15%	3% (†)	
$S(\omega  K_s^0)$	0.17	0.03 (	(*)	$\mathcal{B}(B \to \omega \gamma)$	30%	5%	
$S(f_0K_s^0)$	0.12	0.02 (	(*)	$A_{CP}(B \to K^*\gamma)$	0.007 (†)	0.004 († *)	
				$A_{CP}(B  ightarrow  ho \gamma)$	$\sim 0.20$	0.05	
$\gamma \ (B \to DK, D \to CP \text{ eigenstates})$	$\sim 15^{\circ}$	$2.5^{\circ}$	0	$A_{CP}(b  ightarrow s \gamma)$	$0.012(\dagger)$	0.004 (†)	
$\gamma \ (B \to DK, D \to \text{suppressed states})$		$2.0^{\circ}$		$A_{CP}(b ightarrow (s+d)\gamma)$	0.03	0.006 (†)	
$\gamma \ (B \to DK, D \to \text{multibody states})$	$\sim9^{\circ}$	$1.5^{\circ}$		$S(K^0_s\pi^0\gamma)$	0.15	0.02 (*)	
$\gamma \ (B  o DK,  ext{ combined})$	$\sim 6^{\circ}$	1-2	0	$S( ho^0\gamma)$	possible	0.10	
$\alpha \ (B \to \pi \pi)$	$\sim 16^{\circ}$	3°		$A_{CP}(B  o K^* \ell \ell)$	7%	1%	
$\alpha \; (B  ightarrow  ho  ho)$	$\sim 7^{\circ}$	$1-2^{\circ}$	(*)	$A^{FB}(B \to K^*\ell\ell)s_0$	25%	9%	
$\alpha \; (B  o  ho \pi)$	$\sim 12^{\circ}$	2°		$A^{FB}(B \to X_{s}\ell\ell)s_{0}$	35%	5%	
$\alpha \ (\text{combined})$	$\sim 6^{\circ}$	$1-2^{\circ}$	(*)	$\mathcal{B}(B \to K \nu \overline{\nu})$	visible	20%	
				$\mathcal{B}(B \to \pi \nu \bar{\nu})$ $\mathcal{B}(B \to \pi \nu \bar{\nu})$	4131016	possible	
$2\beta + \gamma \ \left( D^{(*)\pm}\pi^{\mp}, \ D^{\pm}K_{g}^{0}\pi^{\mp} \right)$	20°	5°		$B(B \rightarrow \pi \nu \nu)$		possible	
- Dhusios	Sensitivit	v	B <sub>s</sub> Physics @ Y(5S)				
τ Physics		<u> </u>	Observa	•		Error with $30 \text{ ab}^{-1}$	
${\cal B}( au  o \mu  \gamma)$	$2 \times 10^{-9}$		ΔΓ		$0.16 \text{ ps}^{-1}$	$0.03 \text{ ps}^{-1}$	
			Г		$0.07 \text{ ps}^{-1}$	$0.01 \text{ ps}^{-1}$	
${\cal B}( au  o e  \gamma)$	$2 \times 10^{-9}$			angular analysis	20°	8°	
${\cal B}( au  o \mu  \mu  \mu)$	$2 \times 10^{-10}$	)	$A^s_{ m SL}$		0.006	0.004	
			$A_{ m CH}$		0.004	0.004	
$\mathcal{B}( au  ightarrow eee)$	$2 \times 10^{-10}$	)		$ \mu^+\mu^-)$	-	$< 8 \times 10^{-9}$	
<b>n</b> ( )			$ V_{td}/V_{ts} $		0.08	0.017	
${\cal B}( au  o \mu \eta)$	$4 \times 10^{-10}$	,	$\mathcal{B}(B_s -$		38%	7%	
	$a_{1} = 10^{-10}$		$eta_s$ from	-	$10^{\circ}$	3°	
${\cal B}( au  o e\eta)$	$6 imes10^{-10}$	,	$\beta_s$ from	$B_s \to K^0 \bar{K}^0$	24°	11°	
${\cal B}( au  o \ell K^0_s)$	$2  imes 10^{-10}$	)					
$\sim (r + c m_S)$	2 / 10						

Charm n	nixing	and C	P				
Mode	Observable	e $\Upsilon(4S)$	$\psi(3770)$				
		$(75 \text{ ab}^{-1})$	$(300 \text{ fb}^{-1})$				
$D^0 \rightarrow K^+ \pi^-$	$x'^2$	$3 \times 10^{-5}$					
- 0	y'	$7 \times 10^{-4}$					
$D^0 \to K^+ K^-$ $D^0 \to K^0_S \pi^+ \pi^-$	$y_{CP}$	$5 \times 10^{-4}$					
$D^* \rightarrow K_S^* \pi^+ \pi$	x	$4.9 \times 10^{-4}$ $3.5 \times 10^{-4}$					
	$y \\  q/p $	$3 \times 10^{-2}$					
	$\phi$	2°					
$\psi(3770) \rightarrow D^0 \overline{D}^0$	$x^2$		$(1-2) \times 10^{-5}$				
	y		$(1\!-\!2) \times 10^{-3}$				
	$\cos \delta$		(0.01 - 0.02)				
	ONC						
Charm F	CINC		Sensitivity				
$D^0 \rightarrow e^+ e^-, I$	$\overline{D^0 \to \mu^+ \mu^-}$	,-	$1  imes 10^{-8}$				
$D^0 \to \pi^0 e^+ e^-, D^0 \to \pi^0 \mu^+ \mu^- \qquad 2 \times 10^{-8}$							
$D^0  ightarrow \eta e^+ e^-,  D^0  ightarrow \eta \mu^+ \mu^- \qquad 3 imes 10^-$							
$D^0  ightarrow K^0_s e^+ e^-$	$3 imes 10^{-8}$						
$D^+ \to \pi^+ e^+ e^-,  D^+ \to \pi^+ \mu^+ \mu^- \qquad 1  imes 10^{-8}$							
$D^0  ightarrow e^\pm \mu^\mp$	$1 imes 10^{-8}$						
$D^+  ightarrow \pi^+ e^{\pm} \mu^{\pm}$	$1 imes 10^{-8}$						
$D^0 \to \pi^0 e^{\pm} \mu^{\mp}$	$2 imes 10^{-8}$						
$D^0  o \eta e^{\pm} \mu^{\mp}$	$3 imes 10^{-8}$						
$D^0  ightarrow K^0_s e^\pm \mu^\pm$	$3 imes 10^{-8}$						
$D^+ \rightarrow \pi^- e^+ e^-$	$1 imes 10^{-8}$						
$D^+  o \pi^- \mu^+ \mu$	$1 imes 10^{-8}$						
$D^+ \to \pi^- e^\pm \mu^+$	$1 imes 10^{-8}$						

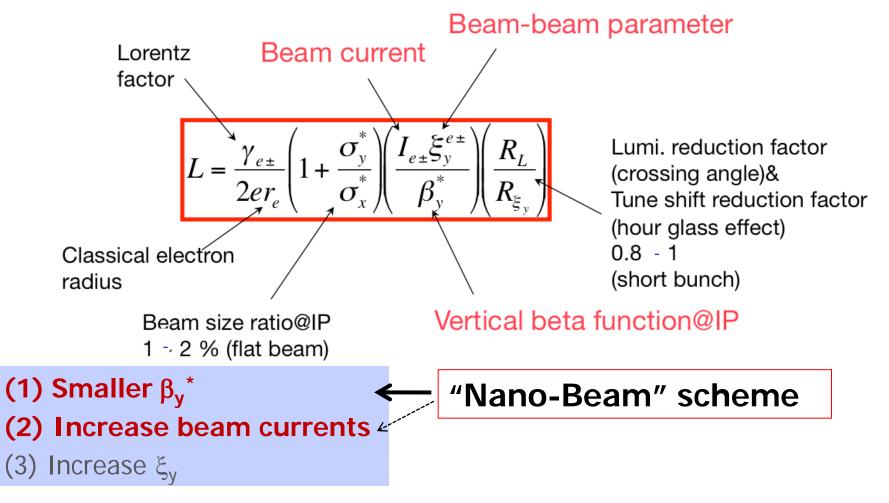
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### Relation between the Super B Factory and the LHC

- Physics motivation is independent of LHC.
  - If LHC finds NP, precision flavour physics is compulsory.
  - If LHC finds no NP, high statistics  $B/\tau$  decays would be a unique way to search for the >TeV scale physics (=TeV scale in case of MFV).

#### How big is a nano-beam?





Collision with very small spot-size beams

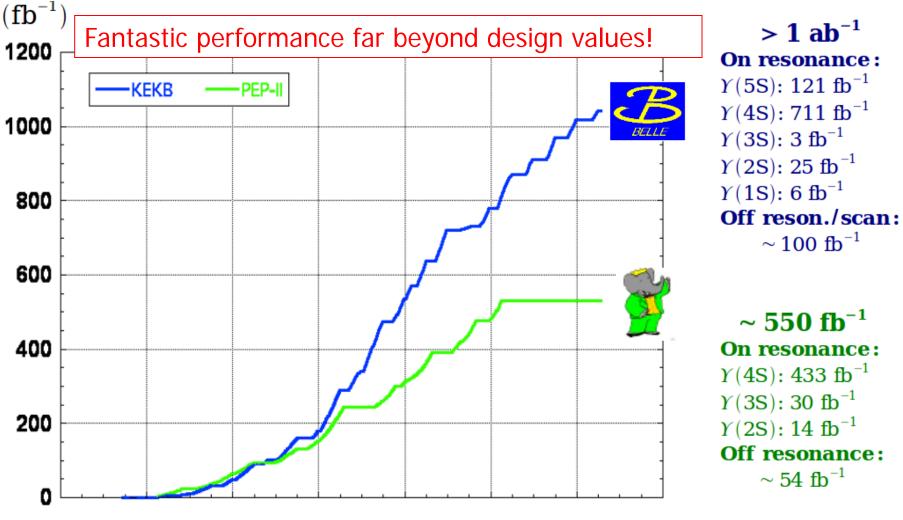
Invented by Pantaleo Raimondi for SuperB

# Belle II International Collaboration



#### 15 countries/regions, ~60 institutions, ~400 collaborators

#### Integrated luminosity at B factories



<sup>1998/1 2000/1 2002/1 2004/1 2006/1 2008/1 2010/1 2012/1</sup>